





THE MICROSCOPE MISSION : TWO YEARS BEFORE THE LAUNCH



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ESA Copernicus programme's Sentinel satellites !





ASAP SOYUZ configuration

- ✓ Copernicus programme : Global Monitoring for Environment (EC with ESA partnership)
- ✓ Five families of Sentinel satellites
- Sentinel-1 is a two satellites constellation with prime objectives of Land and Ocean monitoring :
 C-Band SAR data (accurate imaging in all weathers) following ERS-2 and Envisat.
- ✓ Sentinel 1a ready for launch in spring 2014 : heliosynchronous orbit at 786 km
- ✓ Sentinel 1b to be ready for launch as soon as sept. 2015 with Soyouz (object: end of 2015)



ICROSCOP



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MICROSCOPE UFF TEST RATIONALE



- ✓ Physics is not completely understood → new Physics
 - → New experiments
 - → New type of results
- \checkmark UFF violation \rightarrow one of the invariance of the EEP (UFF, LPI, LLI) violated!

 $\delta_{12} = 2 \frac{\frac{m_{g1}}{m_{i1}} - \frac{m_{g2}}{m_{i2}}}{\frac{m_{g1}}{m_{i1}} + \frac{m_{g2}}{m_{i2}}} \approx \frac{m_{g1}}{m_{i1}} - \frac{m_{g2}}{m_{i2}} = 0? ?$

- ✓ MICROSCOPE Objective : 10⁻¹⁵ accuracy
- ✓ MICROSCOPE is the first accurate UFF test in space
- Scientific results + Return on Space technology limitations
 - Thermal, magnetic, structural, acceleration stabilities @ picometer/s²
 - On board calibration with satellite control
 - Accurate pointing with SST and Angular Accelerometer
 - Scientific Mission Center with Mission Scenario Management



MISSION SCENARIO



- **Reference scenario** is established before the launch : list of sequences (transition or scientific session or technical session):
 - Commissionning step 1: 29 days, operation of all sub-systems & payload verified
 - Commissionning step 2: 20 days, drag free and calibration operation validated
 - Preliminary tests and Performance tests : 25 + 29 days
 - *EP tests : 92 + 52 days*
 - (Calibrations + 2 spins + 2 inertial orientations+ 2 test mass centring) x (EP + REF)
 - Complementary EP tests : 71 days

Breaks periods with satellite in operating mode without thrusters & gas consumption are scheduled and can be added (used to take advantage of the obtained results).

• Working scenario to be executed:

- Cover 1 month
- Is up-dated every weak and validated through Drag-free Expertise and Control Center
- Sequences mentionned as :

 $AE = to be Executed (Q = qualified or AQ = to be qualified) \rightarrow AC = to be confirmed \rightarrow C=confirmed \rightarrow EC = Running \rightarrow E = Execuded or EI = Executed but non successful,$

• Executed scenario updates the reference scenario of the whole mission to compute the whole gas consumption and predict the offered possibilities.







- 2 test masses made of different composition
- Gravitational Source : the Earth
- Kinematic Acceleration : the orbital motion
- Identical initial conditions of motion
- Permanent pico-meter control of the 2 masses
- Measurement = Necessary forces to control the same orbital motion
- No fluctuations of the mass environment due to relative motion
- Centring : 20 µm when the mass are levitated
- \rightarrow Gravity Gradients corrected or centring controlled @ 0.1 µm in orbital plane (X,Z)
- Satellite imposes the common motion : reduced → instrument better operation



Z/μ

0,40286

0.46309

B/μ

1,008009

1.008911

Pt

Ti

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MICROSCOPE PROVIDED MEASUREMENTS



$$\frac{\vec{F} el_{i}}{m_{li}} - \frac{\vec{F} el_{j}}{m_{lj}} = \left(\delta_{j} - \delta_{i}\right)\vec{g}\left(O_{j}\right) + \left(1 + \delta_{i}\right)\left[T\right]\overrightarrow{O_{i}O_{j}} - R_{ln,COR}\left(\overrightarrow{O_{i}O_{j}}\right) - \frac{\vec{F} pa_{i}}{m_{li}} + \frac{\vec{F} pa_{j}}{m_{lj}}$$



• centering

• shape : spherical inertia, multipoles

material density homogeneity

$$R_{In,COR}\left(\overrightarrow{O_{sat}O_{k}}\right) = \overrightarrow{\Omega} \wedge \overrightarrow{O_{sat}O_{k}} + \overrightarrow{\Omega} \wedge (\overrightarrow{\Omega} \wedge \overrightarrow{O_{sat}O_{k}}) + 2[\Omega]\overrightarrow{O_{sat}O_{k}} + \overrightarrow{O_{sat}O_{k}}$$

Angular acceleration & centrifugal acceleration : to be controlled
Coriolis & Cinematic relative acceleration

→ stability of the ULE configuration and electrostatic servo-conrol



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2 years mission duration : fine survey of gas consumption Heliosynchronous orbit ~ 710 km \rightarrow 1,7 10⁻⁴ Hz Passive temperature stabilities

Compensation of the drag by GAIA type thrusters Attitude control without gyro. and wheels Inertial and rotating pointing \rightarrow 1 mHz

No moving masses and structural motions @ $\rm f_{EP}$ Position and attitude sufficiently well known

Payload contributes to s/c motion control S/C contributes to Payload outputs



2 differential electrostatic accelerometers in thermal cocoon





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magnetic cocoon



payload at the center of the satellite : -for thermal stability -for spin mode -for self gravity



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MICROSCOP





MCROSCOPE

NCROSCOPE



MICROSCOPE Space Lab. with 6 DoF Controlled to the benefit of the environment stability



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THE FRENCH AEROSPACE LAB

Space Electrostatic accelerometers for Earth gravity field recovery

GRACE (NASA-JPL), March 2002 – 2015 ?





GOCE (ESA), March 2009 – October 2013 ?



- Γn: 2.0·10⁻¹² ms⁻² /Hz^{1/2} • Γmax: 610⁻⁶ ms⁻²
- [5·10⁻³; 10⁻¹] Hz





Courtesy TAS-F





Space Electrostatic accelerometers for Earth gravity field recovery



GRACE (NASA-JPL), March 2002 - 2015 ?



altitude~500km • Fn: 1.0.10⁻¹⁰ ms⁻² /Hz^{1/2} • Fmax: 510⁻⁵ ms⁻² • [0.1.10⁻³; 10⁻¹] Hz

Today : 3971 days in orbit

GOCE (ESA), March 2009 – October 2013 ?

altitude~260km

• Fmax: 610⁻⁶ ms⁻² • [5.10⁻³; 10⁻¹] Hz



and individual diagonal gravity gradient (red, ormance for calibrated gradiometer (with nonplied). Note that <30 mHz the noise is not visible gradient curves, since signal dominates

• Γn: 2.0-10⁻¹² ms⁻² /Hz^{1/2} ASH_{1.4}: 3.9 10⁻¹² m/s²/Hz^{1/2} ASH_{2.5}: 3.1 10⁻¹² m/s²/Hz^{1/2}

ASH_{3.6}: 6.7 10⁻¹² m/s²/Hz^{1/2}

GOCE	MICROSCOPE
Gold wire : Ø= 5 μm	Ø=7.5 μm
PT-Rh Proof mass : m= 320g	m=1400 - 307 g
Gap Y,Z: e = 299 μm	e= 600 μm
PM Polarisation : Vp = 7.5 V	Vp =5 V
Detection : Vd = 7.6 V @ 100 KHz	Vd = 7.07 V @ 100 KHz
Detector gain 1.7 mV / nano-m	0.3 - 0.26 mV / nano-m
Scale factor :	
Science data 1. 10 ⁻⁷ ms ⁻² /V	1.8 - 2.1 10 ⁻⁷ ms ⁻² /V
DFACS data 17. 10 ⁻⁶ ms ⁻² /V	0.7- 1.7 10 ⁻⁶ ms ⁻² /V
Range ± 6.5 10 ⁻⁶ ms ⁻²	± 4.8 - 4.6 10 ⁻⁷ ms ⁻²
Expected Res. < 2 10 ⁻¹² ms ⁻² Hz ^{-1/2}	< 2 10 ⁻¹² ms ⁻² Hz ^{-1/2}

ONERA THE FRENCH AEROSPACE LAB



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Courtesy TAS-F

MICROSCOPE : A dedicated instrument





Instrument Design





Instrument Design

Instrument Design

MICROSCOPE

MICROSCOPE

Reference plate with silica top hat and fingers

Reference plate with silica top hat and fingers

Reference plate with silica top hat and fingers

Reference plate with silica top hat and fingers

Reference plate with inner radial electrode rod

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INTEGRATION FM 2 : Platinum / Titanium (3/4)

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FEEU and ICU qualification

X inner: 2.6 10^5 V/m \rightarrow <10µV.Hz^{1/2} = 3.8 10^{-11} m.Hz^{1/2}

Axe	Résultat [µV/℃]	Spécification
X1i	64.68	<300µV/C
X1e	16.66	<300µV/C
Y1i	11.70	<20µV/C
Y1e	14.68	<20µV/C
Y2i	8.40	<20µV/C
Y2e	9.34	<20µV/C
X2i	36.56	<300µV/C
X2e	130.93	<1800µV/C
Z1i	12.15	<20µV/C
Z1e	37.63	<20µV/C
Z2i	16.37	<20µV/C
Z2e	13.26	<20µV/C

X outer: $3.0\ 10^5\ V/m \rightarrow < 6\mu V.Hz^{1/2} = 2.0\ 10^{-11}\ m.Hz^{1/2}$ Y,Z inner: $2.3\ 10^5\ V/m \rightarrow < 6\mu V.Hz^{1/2} = 2.6\ 10^{-11}m.Hz^{1/2}$ Y,Z outer: $3.1\ 10^5\ V/m \rightarrow < 4\mu V.Hz^{1/2} = 1.3\ 10^{-11}m.Hz^{1/2}$ The 3 FEEU FM, successfully tested in performance :

- noise + bias + linearity + bandwidth + thermal sensitivity
- Interface with FM ICUs

Now, under potting after increase of the read-out range, Qualification under vibrations and thermal cycles are scheduled in Dec. 2013.

FM Interface Control Unit :DSP Hardware

1 DSP board per differential accelerometer (No redundancy) : 1 DSP → 1 SU → 12 servo-loops channels

1 DSP = 1 Oslink customer

Architecture on TSC21020F:

•Rad-tolerant FPGA

•SEL immune SRAM (SEU protected by EDAC)

Software and Hardware tested, accuracy verified.

•PROM containing the master (Boot) software (IMSW)

•EEPROM containing the application software (ASW) and the parameter tables.

isgend Implemented in the DSP ormpanion chip (FPGA) Use the intervent of the DSP ormpanion chip (FPGA) Use the intervent of the DSP ormpanion chip (FPGA) Use the intervent of the DSP ormpanion chip (FPGA) DSP: IRQ Timer Use the intervent of the DSP ormpanion chip (FPGA) DATA SRAM (IdB + check) ToFrom FEEU PROA ToFrom FEEU Controller Cock Cock

- DSP hardware now compliant (more robust chronogram) :
 - > With the whole range of operating temperature
 - > And with the 2 years duration of the mission
- Tests have been successfully performed
- Software 2.6 to be delivered at end of November.

Instrument status and performance verification

Sensors:

Qualification, now performed with demonstration of resistance to launch vibrations, chocks, aging (gold wire); FM 2, integrated and under tests; FM 1, integration running;

To be delivered in March 2014

<u>Analog electronics:</u>
 FM Tested and being potted after full range adjustment;
 To be delivered in Feb. 2014

<u>Digital electronics:</u>
 Robustness to increase of temperature now insured;
 Software to be up-dated;
 To be delivered in Feb. 2014

Documentation: In progress

Error budget

Now performed with QM actual values and satellite expected environment Spin mode : 1,12 10⁻¹⁵ over 20 orbits and 0,66 10⁻¹⁵ over 120 orbits Inertial mode : 1,42 10⁻¹⁵ over 120 orbits Both limited by the sensor noise, the SU gradients of temperature variations, the SU and FEEU temperature variations.

Instrument status and performance verification

Performance test session : 29 days

- ✓ Verification of acceleration output linearity
- ✓ Sensibility of output linearity to static TM position (along the 3 axes)
- ✓ Variation of the electrostatic configuration
 - ✓ through test mass DC potential
 - → observation of bias and noise
 - ✓ through test-mass sine motion : change of geometry
 - ✓ through S/C sine motion : change of electrode voltages
- ✓ Evaluation of couplings and TM self gravity
- ✓ Evaluation of Magnetic sensitivity through magneto-torquer actuations
- ✓ Evaluation of thermal sensitivity of SU and FEEU with dedicated thermistances

Calibration : 3 phases of 14 days

- ✓ Before and after
- ✓ EP and REF

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bns, chocks, aging (gold wire);

CALIBRATION : 2 servo-loops to generate well known acceleration outputs

✓ Drag compensation loop \rightarrow To excite the linear satellite motion

- \rightarrow Common excitation \rightarrow Differential outputs vs drag-free point
- ✓ Attitude S/C control through SST <u>and</u> angular accelerometer
 - \rightarrow To oscillate the S/C \rightarrow Differential outputs vs eccentricity or instrument attitude vs SST
- ✓ Proof- mass oscillation \rightarrow Elect. Conf. modif. Or Corriolis effects

Operational & scientific organization

NICROSCOPE

3 levels

The MICROSCOPE Science Working Group promotes the exploitation of the data & is responsible in particular for:

- Supervising and approving the evaluation and the validation of the performance
- Approving the final scientific data products to be distributed to the community,
- Promoting the exploitation of the data and the diffusion of the information (colloquia...).

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MICROSCOPE **Thanks to MICROSCOPE present partners** with Cnes Courtesy esa AZU cnes Observatoire de la Côte d'Azur **CENTRE NATIONAL D'ÉTUDES SPATIALES** ONERA Non Calibrated 1E-15 EP signal @ 730km THE FRENCH AEROSPACE LAB DLR PB **Physikalisch-**ZVYW **Technische Bundesanstalt** ONERA

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THANK YOU FOR YOUR ATTENTION

QUESTIONS ?

